
Cold Diffusion: Inverting Arbitrary Image Transforms Without Noise

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Abstract

1 Standard diffusion models involve an image transform – adding Gaussian noise –
2 and an image restoration operator that inverts this degradation. We observe that the
3 generative behavior of diffusion models is not strongly dependent on the choice
4 of image degradation, and in fact an entire family of generative models can be
5 constructed by varying this choice. Even when using completely deterministic
6 degradations (e.g., blur, masking, and more), the training and test-time update rules
7 that underlie diffusion models can be easily generalized to create generative models.
8 The success of these fully deterministic models calls into question the community’s
9 understanding of diffusion models, which relies on noise in either gradient Langevin
10 dynamics or variational inference, and paves the way for generalized diffusion
11 models that invert arbitrary processes.

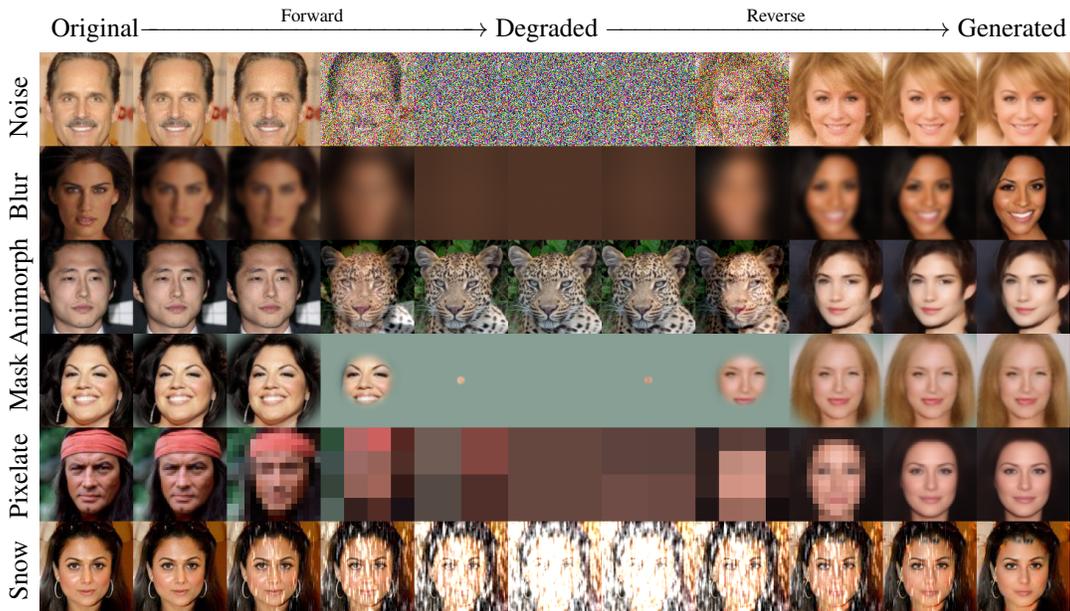


Figure 1: Demonstration of the forward and backward processes for both hot and cold diffusions. While standard diffusions are built on Gaussian noise (top row), we show that generative models can be built on arbitrary and even noiseless/cold image transforms, including the ImageNet-C *snowification* operator, and an *animorphosis* operator that adds a random animal image from AFHQ.

12 **1 Introduction**

13 Diffusion models have recently emerged as powerful tools for generative modeling [Ramesh et al.,
14 2022]. Diffusion models come in many flavors, but all are built around the concept of random noise
15 removal; one trains an image restoration/denoising network that accepts an image contaminated with
16 Gaussian noise, and outputs a denoised image. At test time, the denoising network is used to convert
17 pure Gaussian noise into a photo-realistic image using an update rule that alternates between applying
18 the denoiser and adding Gaussian noise. When the right sequence of updates is applied, complex
19 generative behavior is observed.

20 The origins of diffusion models, and also our theoretical understanding of these models, are strongly
21 based on the role played by Gaussian noise during training and generation. Diffusion has been
22 understood as a random walk around the image density function using *Langevin dynamics* [Sohl-
23 Dickstein et al., 2015, Song and Ermon, 2019], which requires Gaussian noise in each step. The walk
24 begins in a high temperature (heavy noise) state, and slowly anneals into a “cold” state with little if
25 any noise. Another line of work derives the loss for the denoising network using variational inference
26 with a Gaussian prior [Ho et al., 2020, Song et al., 2021a, Nichol and Dhariwal, 2021].

27 In this work, we examine the need for Gaussian noise, or any randomness at all, for diffusion models
28 to work in practice. We consider *generalized diffusion models* that live outside the confines of the
29 theoretical frameworks from which diffusion models arose. Rather than limit ourselves to models
30 built around Gaussian noise, we consider models built around arbitrary image transformations like
31 blurring, downsampling, etc. We train a restoration network to invert these deformations using a
32 simple ℓ_p loss. When we apply a sequence of updates at test time that alternate between the image
33 restoration model and the image degradation operation, generative behavior emerges, and we obtain
34 photo-realistic images.

35 The existence of *cold diffusions* that require no Gaussian noise (or any randomness) during training
36 or testing raises questions about the limits of our theoretical understanding of diffusion models. It
37 also unlocks the door for potentially new types of generative models with very different properties
38 than conventional diffusion seen so far.

39 **2 Background**

40 Both the Langevin dynamics and variational inference interpretations of diffusion models rely on
41 properties of the Gaussian noise used in the training and sampling pipelines. From the score-matching
42 generative networks perspective [Song and Ermon, 2019, Song et al., 2021b], noise in the training
43 process is critically thought to expand the support of the low-dimensional training distribution to
44 a set of full measure in ambient space. The noise is also thought to act as data augmentation to
45 improve score predictions in low density regions, allowing for mode mixing in the stochastic gradient
46 Langevin dynamics (SGLD) sampling. The gradient signal in low-density regions can be further
47 improved during sampling by injecting large magnitudes of noise in the early steps of SGLD and
48 gradually reducing this noise in later stages.

49 Kingma et al. [2021] propose a method to learn a noise schedule that leads to faster optimization.
50 Using a classic statistical result, Kadkhodaie and Simoncelli [2021] show the connection between
51 removing additive Gaussian noise and the gradient of the log of the noisy signal density in determin-
52 istic linear inverse problems. Here, we shed light on the role of noise in diffusion models through
53 theoretical and empirical results in applications to inverse problems and image generation.

54 Iterative neural models have been used for various inverse problems [Romano et al., 2016, Metzler
55 et al., 2017]. Recently, diffusion models have been applied to them [Song et al., 2021b] for the
56 problems of deblurring, denoising, super-resolution, and compressive sensing [Whang et al., 2021,
57 Kawar et al., 2021, Saharia et al., 2021, Kadkhodaie and Simoncelli, 2021].

58 Although not their focus, previous works on diffusion models have included experiments with
59 deterministic image generation [Song et al., 2021a, Dhariwal and Nichol, 2021, Karras et al.,
60 2022] and in selected inverse problems [Kawar et al., 2022]. Recently, Rissanen et al. [2022] use a
61 combination of Gaussian noise and blurring as a forward process for diffusion. Though they show
62 the feasibility of a different degradation, here we show definitively that noise is not a *necessity* in
63 diffusion models, and we observe the effects of removing noise for a number of inverse problems.

64 Despite prolific work on generative models in recent years, methods to probe the properties of learned
 65 distributions and measure how closely they approximate the real training data are by no means closed
 66 fields of investigation.

67 Indirect feature space similarity metrics such as Inception Score [Salimans et al., 2016], Mode Score
 68 [Che et al., 2016], Frechet inception distance (FID) [Heusel et al., 2017], and Kernel inception
 69 distance (KID) [Bińkowski et al., 2018] have been proposed and adopted to some extent, but they
 70 have notable limitations [Barratt and Sharma, 2018]. To adopt a popular frame of reference, we will
 71 use FID as the feature similarity metric for our experiments.

72 3 Generalized Diffusion

73 Standard diffusion models are built around two components. First, there is an image degradation
 74 operator that contaminates images with Gaussian noise. Second, a trained restoration operator is
 75 created to perform denoising. The image generation process alternates between the application of
 76 these two operators. In this work, we consider the construction of generalized diffusions built around
 77 arbitrary degradation operations. These degradations can be randomized (as in the case of standard
 78 diffusion) or deterministic.

79 3.1 Model components and training

80 Given an image $x_0 \in \mathbb{R}^N$, consider the *degradation* of x_0 by operator D with severity t , denoted
 81 $x_t = D(x_0, t)$. The output distribution $D(x_0, t)$ of the degradation should vary continuously in t ,
 82 and the operator should satisfy $D(x_0, 0) = x_0$.

83 In the standard diffusion framework, D adds Gaussian noise with variance proportional to t . In our
 84 generalized formulation, we choose D to perform various other transformations such as blurring,
 85 masking out pixels, downsampling, and more, with severity that depends on t . We explore a range of
 86 choices for D in Section 4.

87 We also require a *restoration* operator R that (approximately) inverts D . This operator has the property
 88 that $R(x_t, t) \approx x_0$. In practice, this operator is implemented via a neural network parameterized by
 89 θ . The restoration network is trained via the minimization problem

$$\min_{\theta} \mathbb{E}_{x \sim \mathcal{X}} \|R_{\theta}(D(x, t), t) - x\|,$$

90 where x denotes a random image sampled from distribution \mathcal{X} and $\|\cdot\|$ denotes a norm, which we
 91 take to be ℓ_1 in our experiments. We have so far used the subscript R_{θ} to emphasize the dependence
 92 of R on θ during training, but we will omit this symbol for simplicity in the discussion below.

93 3.2 Sampling from the model

94 After choosing a degradation D and training
 95 a model R to perform the restoration,
 96 these operators can be used in tandem to in-
 97 vert severe degradations by using standard
 98 methods borrowed from the diffusion liter-
 99 ature. For small degradations ($t \approx 0$), a
 100 single application of R can be used to ob-
 101 tain a restored image in one shot. However,
 102 because R is typically trained using a sim-
 103 ple convex loss, it yields blurry results when
 104 used with large t . Rather, diffusion models
 105 [Song et al., 2021a, Ho et al., 2020] per-
 106 form generation by iteratively applying the
 107 denoising operator and then adding noise
 108 back to the image, with the level of added
 109 noise decreasing over time. This is the standard update sequence in Algorithm 1.

Algorithm 1 Naive Sampling

Input: A degraded sample x_t
for $s = t, t - 1, \dots, 1$ **do**
 $\hat{x}_0 \leftarrow R(x_s, s)$
 $x_{s-1} = D(\hat{x}_0, s - 1)$
end for
Return: x_0

Algorithm 2 Transformation Agnostic Cold Sampling

Input: A degraded sample x_t
for $s = t, t - 1, \dots, 1$ **do**
 $\hat{x}_0 \leftarrow R(x_s, s)$
 $x_{s-1} = x_s - D(\hat{x}_0, s) + D(\hat{x}_0, s - 1)$
end for

110 When the restoration operator is perfect, *i.e.* when $R(D(x_0, t), t) = x_0$ for all t , one can easily see
 111 that Algorithm 1 produces exact iterates of the form $x_s = D(x_0, s)$. But what happens for imperfect

112 restoration operators? In this case, errors can cause the iterates x_s to wander away from $D(x_0, s)$,
 113 and inaccurate reconstruction may occur.

114 We find that the standard sampling approach in Algorithm 1 (explained further in A.8) works well
 115 for noise-based diffusion, possibly because the restoration operator R has been trained to correct
 116 (random Gaussian) errors in its inputs. However, we find that it yields poor results in the case of cold
 117 diffusions with smooth/differentiable degradations as demonstrated for a deblurring model in Figure
 118 2. We propose Transformation Agnostic Cold Sampling (TACoS) in Algorithm 2, which we find to
 119 be superior for inverting smooth, cold degradations.

120 This sampler has important mathematical properties that enable it to recover high quality results.
 121 Specifically, for a class of linear degradation operations, it can be shown to produce exact reconstruction
 122 (i.e. $x_s = D(x_0, s)$) even when the restoration operator R fails to perfectly invert D . We discuss
 123 this in the following section.

124 3.3 Properties of TACoS

125 It is clear from inspection that both Algo-
 126 rithms 1 and 2 perfectly reconstruct the it-
 127 erate $x_s = D(x_0, s)$ for all $s < t$ if the
 128 restoration operator is a perfect inverse for
 129 the degradation operator. Hence in this sec-
 130 tion, we will discuss the reconstruction op-
 131 erator that fails to reconstruct the image per-
 132 fectly i.e. incurs error. We first analyze the
 133 stability of these algorithms to errors in the
 134 restoration operator and then theoretically
 135 show that for a simple blur degradation, the
 136 error incurred using algorithm 1 is always
 137 greater than algorithm 2.

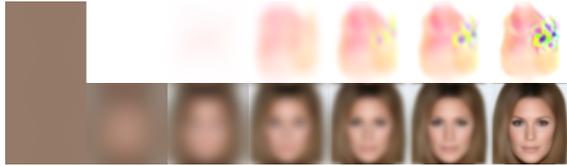


Figure 2: Comparison of sampling methods for unconditional generation using cold diffusion on the CelebA dataset. Iterations 2, 4, 8, 16, 32, 64, 128, 192, and 256 are presented. **Top:** Algorithm 1 produces compounding artifacts and fails to generate a new image. **Bottom:** TACoS succeeds in sampling a high quality image without noise.

138 For small values of x and s , TACoS as described in 2 is tolerant of error in the restoration operator R .
 139 To see why, consider a problem with linear degradation function of the form $D(x, s) \approx x + s \cdot e$ for
 140 a constant vector e . We choose this ansatz because the Taylor expansion of any smooth degradation
 141 $D(x, s)$ around $x = x_0, s = 0$ has the form $D(x, s) \approx x + s \cdot e(x) + \text{HOT}$ where HOT denotes
 142 higher order terms. Note, however, the analysis below requires e to be a constant that does not depend
 143 on x . The constant/zeroth-order term in this Taylor expansion is zero because we assumed above that
 144 the degradation operator satisfies $D(x, 0) = x$.

145 For a degradation $D(x, s)$ and any restoration operator R , the term x_{s-1} in TACoS becomes

$$\begin{aligned} x_s - D(R(x_s, s), s) + D(R(x_s, s), s - 1) &= D(x_0, s) - D(R(x_s, s), s) + D(R(x_s, s), s - 1) \\ &= x_0 + s \cdot e - R(x_s, s) - s \cdot e + R(x_s, s) + (s - 1) \cdot e = x_0 + (s - 1) \cdot e = D(x_0, s - 1) \end{aligned}$$

146 By induction, we see that the algorithm outputs the value $x_s = D(x_0, s)$ for all $s < t$, regardless of
 147 the choice of R . In other words, for *any* choice of R , the iteration behaves the same as it would when
 148 R is a perfect inverse for the degradation D .

149 By contrast, Algorithm 1 does not enjoy this behavior even for small values of s . In fact, when
 150 R is not a perfect inverse for D , x_0 is not a fixed point of the update rule in Algorithm 1 because
 151 $x_0 \neq D(R(x_0, 0), 0) = R(x_0, 0)$ and hence errors compound. If R does not perfectly invert D , we
 152 should expect Algorithm 1 to incur errors, even for small values of s . Meanwhile, for small values of
 153 s , the behavior of D approaches its first-order Taylor expansion, and Algorithm 2 becomes immune
 154 to errors in R . Figure 2 demonstrates the stability of TACoS described in Algorithm 2 vs. Algorithm
 155 1 for a deblurring model.

156 The above analysis is not a complete convergence theory but rather highlights a desirable theoretical
 157 property of our method that a naive sampler lacks. However, we can prove that for a *toy* problem
 158 in which the blur operator removes one frequency at a time, the error incurred by sampling using
 159 Algorithm 1 is greater than the error incurred from using Algorithm 2. We present the proof of this
 160 claim in A.9.

161 **4 Generalized Diffusions with Various Transformations**

162 In this section, we take the first step towards cold diffusion by reversing different degradations and
 163 hence performing conditional generation. We will extend our methods to perform unconditional
 164 (i.e. from scratch) generation in Section 5. We empirically evaluate generalized diffusion models
 165 trained on different degradations with TACoS proposed in Algorithm 2. We perform experiments
 166 on the vision tasks of deblurring, inpainting, and super-resolution. We perform our experiments on
 167 MNIST [LeCun et al., 1998], CIFAR-10 [Krizhevsky, 2009], and CelebA [Liu et al., 2015]. In each of
 168 these tasks, we gradually remove the information from the clean image, creating a sequence of images
 169 such that $D(x_0, t)$ retains less information than $D(x_0, t - 1)$. For these different tasks, we present
 170 both qualitative and quantitative results on a held-out testing dataset and demonstrate the importance
 171 of the sampling technique described in Algorithm 2. For all quantitative results in this section, the
 172 Frechet inception distance (FID) scores [Heusel et al., 2017] for degraded and reconstructed images
 173 are measured with respect to the testing data. Additional information about the quantitative results,
 174 convergence criteria, hyperparameters, and architecture of the models presented below can be found
 175 in the appendix.

176 **4.1 Deblurring**

177 We consider a generalized diffusion based on a Gaussian blur operation (as opposed to Gaussian
 178 noise) in which an image at step t has more blur than at $t - 1$. The forward process given the Gaussian
 179 kernels $\{G_s\}$ and the image x_{t-1} at step $t - 1$ can thus be written as

$$x_t = G_t * x_{t-1} = G_t * \dots * G_1 * x_0 = \bar{G}_t * x_0 = D(x_0, t),$$

180 where $*$ denotes the convolution operator, which blurs an image using a kernel.

181 We train a deblurring model by minimizing the loss (1), and then use TACoS to invert this blurred
 182 diffusion process for which we trained a DNN to predict the clean image \hat{x}_0 . Qualitative results
 183 are shown in Figure 3 and quantitative results in Table 1. Qualitatively, we can see that images
 184 created using the sampling process are sharper and in some cases completely different as compared
 185 to the direct reconstruction of the clean image. Quantitatively we can see that the reconstruction
 186 metrics such as RMSE and PSNR get worse when we use the sampling process, but on the other hand
 187 FID with respect to held-out test data improves. The qualitative improvements and decrease in FID
 188 show the benefits of the generalized sampling routine, which brings the learned distribution closer
 189 to the true data manifold. Note: we compare the images reconstructed via Algorithm 2, with direct
 190 generation as compared to Algorithm 1. This is because the image reconstruction via Algorithm 1 is
 191 much worse than both direct generation and Algorithm 2. Nevertheless, to back our claim, we present
 192 their results in A.10.

193 In the case of blur operator, the sampling routine can be thought of adding frequencies at each step.
 194 This is because the sampling routine involves the term $D(\hat{x}_0, t) - D(\hat{x}_0, t - 1)$ which in the case
 195 of blur becomes $\bar{G}_t * x_0 - \bar{G}_{t-1} * x_0$. This results in a difference of Gaussians, which is a band
 196 pass filter and contains frequencies that were removed at step t . Thus, in the sampling process, we
 197 sequentially add the frequencies that were removed during the degradation process.

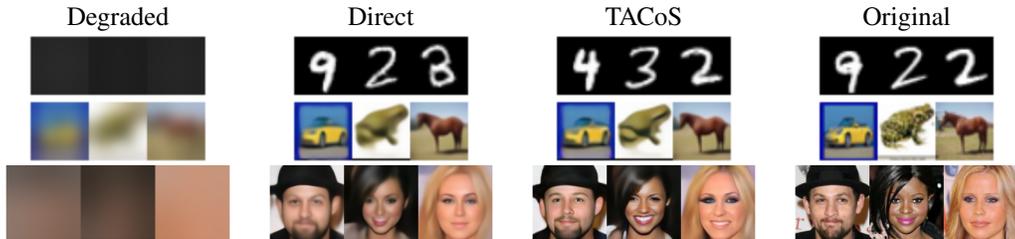


Figure 3: Deblurring models trained on the MNIST, CIFAR-10, and CelebA datasets. **Left to right:** degraded inputs $D(x_0, T)$, direct reconstruction $R(D(x_0, T))$, sampled reconstruction with TACoS described in Algorithm 2, and original image.

Table 1: Quantitative metrics for quality of image reconstruction using deblurring models.

Dataset	Degraded			Sampled			Direct		
	FID	SSIM	RMSE	FID	SSIM	RMSE	FID	SSIM	RMSE
MNIST	438.59	0.287	0.287	4.69	0.718	0.154	5.10	0.757	0.142
CIFAR-10	298.60	0.315	0.136	80.08	0.773	0.075	83.69	0.775	0.071
CelebA	382.81	0.254	0.193	26.14	0.568	0.093	36.37	0.607	0.083

198 **4.2 Inpainting**

199 We define a schedule of transforms that progressively grays-out pixels from the input image. We
 200 remove pixels using a Gaussian mask as follows: For input images of size $n \times n$ we start with a 2D
 201 Gaussian curve of variance β , discretized into an $n \times n$ array. We normalize so the peak of the curve
 202 has value 1, and subtract the result from 1 so the center of the mask as value 0. We randomize the
 203 location of the Gaussian mask for MNIST and CIFAR-10, but keep it centered for CelebA. We denote
 204 the final mask by z_β .

205 Input images x_0 are iteratively masked for T steps via multiplication with a sequence of masks $\{z_{\beta_i}\}$
 206 with increasing β_i . We can control the amount of information removed at each step by tuning the
 207 β_i parameter. In the language of Section 3, $D(x_0, t) = x_0 \cdot \prod_{i=1}^t z_{\beta_i}$, where the operator \cdot denotes
 208 entry-wise multiplication.

209 Figure 4 presents results on test images and compares the output of the inpainting model to the
 210 original image. The reconstructed images display reconstructed features qualitatively consistent
 211 with the context provided by the unperturbed regions of the image. We quantitatively assess the
 212 effectiveness of the inpainting models on each of the datasets by comparing distributional similarity
 213 metrics before and after the reconstruction. Our results are summarized in Table 2. Note, the FID
 214 scores here are computed with respect to the held-out validation set.

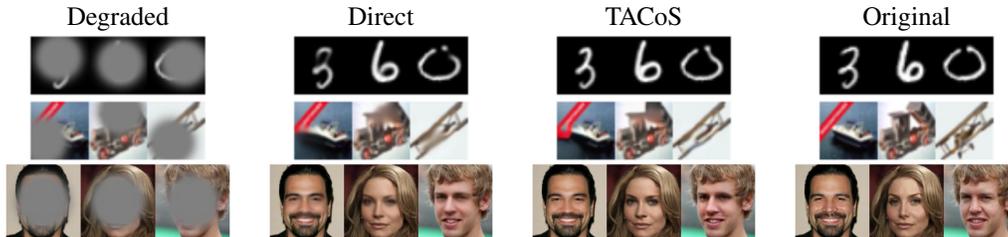


Figure 4: Inpainting models trained on the MNIST, CIFAR-10, and CelebA datasets. **Left to right:** Degraded inputs $D(x_0, T)$, direct reconstruction $R(D(x_0, T))$, sampled reconstruction with TACoS described in Algorithm 2, and original image.

Table 2: Quantitative metrics for quality of image reconstruction using inpainting models.

Dataset	Degraded			Sampled			Direct		
	FID	SSIM	RMSE	FID	SSIM	RMSE	FID	SSIM	RMSE
MNIST	108.48	0.490	0.262	1.61	0.941	0.068	2.24	0.948	0.060
CIFAR-10	40.83	0.615	0.143	8.92	0.859	0.068	9.97	0.869	0.063
CelebA	127.85	0.663	0.155	5.73	0.917	0.043	7.74	0.922	0.039

215 **4.3 Super-Resolution**

216 For this task, the degradation operator downsamples the image by a factor of two in each direction.
 217 The final resolution of x_T is 4×4 for MNIST and CIFAR-10 and 2×2 in the case of Celeb-A.
 218 After each down-sampling, the lower-resolution image is resized to the original image size, using
 219 nearest-neighbor interpolation. More details are available in Appendix A.3

220 Figure 5 presents example testing data inputs for all datasets and compares the output of the super-
 221 resolution model to the original image. Though the reconstructed images are not perfect for the

222 more challenging datasets, the reconstructed features are qualitatively consistent with the context
 223 provided by the low resolution image. Table 3 compares the distributional similarity metrics between
 224 degraded/reconstructed images and test samples.

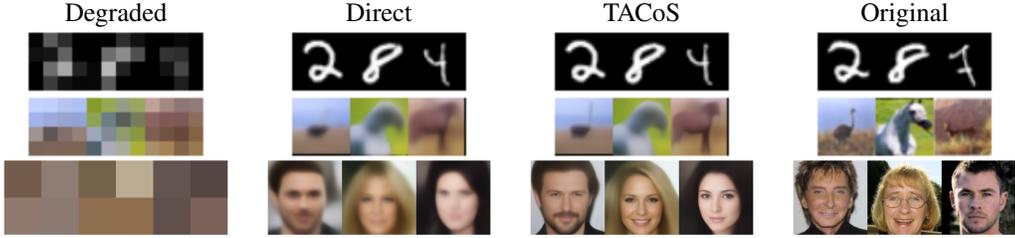


Figure 5: Superresolution models trained on the MNIST, CIFAR-10, and CelebA datasets. **Left to right:** degraded inputs $D(x_0, T)$, direct reconstruction $R(D(x_0, T))$, sampled reconstruction with TACoS described in Algorithm 2, and original image.

Table 3: Quantitative metrics for quality of image reconstruction using super-resolution models.

Dataset	Degraded			Sampled			Direct		
	FID	SSIM	RMSE	FID	SSIM	RMSE	FID	SSIM	RMSE
MNIST	368.56	0.178	0.231	4.33	0.820	0.115	4.05	0.823	0.114
CIFAR-10	358.99	0.279	0.146	152.76	0.411	0.155	169.94	0.420	0.152
CelebA	349.85	0.335	0.225	96.92	0.381	0.201	112.84	0.400	0.196

225 5 Cold Generation

226 Diffusion models can successfully learn the underlying distribution of training data, and thus generate
 227 diverse, high quality images [Song et al., 2021a, Dhariwal and Nichol, 2021, Jolicoeur-Martineau
 228 et al., 2021, Ho et al., 2022]. We will first discuss deterministic generation using Gaussian noise
 229 and then discuss in detail unconditional generation using deblurring. Finally, we provide a proof of
 230 concept that the TACoS described in Algorithm 2 can be extended to other degradations.

231 5.1 Generation using deterministic noise degradation

232 Here we discuss image generation using a noise-based degradation presented in our notation from
 233 Section 3, which we will later prove is equivalent to DDIM [Song et al., 2021a]. We use the following
 234 degradation operator: $D(x, t) = \sqrt{\alpha_t}x + \sqrt{1 - \alpha_t}z$.

235 D is an interpolation between the data point x and a sampled noise pattern $z \in \mathcal{N}(0, 1)$. During
 236 training, D is applied *once* and thus z is sampled once for every image in every batch. However,
 237 sampling involves iterative applications of the degradation operator D , which poses the question of
 238 how to pick z for the sequence of degradations D applied in a single image generation.

239 There are three possible choices for z . The first would be to resample z for each application of D , but
 240 this would make the sampling process nondeterministic for a fixed starting point. Another option is
 241 to sample a noise pattern z *once* for each *separate* image generation and reuse it in each application
 242 of D . In Table 4 we refer to this approach as *Fixed Noise*. Finally, one can calculate the noise vector
 243 z to be used in step t of reconstruction by using the formula

$$\hat{z}(x_t, t) = \frac{x_t - \sqrt{\alpha_t}R(x_t, t)}{\sqrt{1 - \alpha_t}}.$$

244 This method denoted *Estimated Noise* in Table 4 turns out to be equivalent to the deterministic
 245 sampling proposed in Song et al. [2021a]. We discuss this equivalence in detail in Appendix A.6.

246 5.2 Image generation using blur

247 The forward diffusion process in noise-based diffusion models has the advantage that the degraded
 248 image distribution at the final step T is simply an isotropic Gaussian. One can therefore perform

249 (unconditional) generation by first drawing a sample from the isotropic Gaussian, and sequentially
 250 denoising it with backward diffusion.

251 When using blur as a degradation, the fully degraded images do not form a nice closed-form
 252 distribution that we can sample from. They do, however, form a simple enough distribution that can
 253 be modeled with simple methods. Note that every image x_0 degenerates to an x_T that is constant (i.e.,
 254 every pixel is the same color) for large T . Furthermore, the constant value is exactly the channel-wise
 255 mean of the RGB image x_0 , and can be represented with a 3-vector. This 3-dimensional distribution
 256 is easily represented using a Gaussian mixture model (GMM). This GMM can be sampled to produce
 257 the random pixel values of a severely blurred image, which can be deblurred using cold diffusion to
 258 create a new image.

259 Our generative model uses a blurring schedule where we progressively blur each image with a
 260 Gaussian kernel of size 27×27 over 300 steps. The standard deviation of the kernel starts at 1 and
 261 increases exponentially at the rate of 0.01. We then fit a simple GMM with one component to the
 262 distribution of channel-wise means. To generate an image from scratch, we sample the channel-wise
 263 mean from the GMM, expand the 3D vector into a 128×128 image with three channels, and then
 264 apply TACoS.

265 Empirically, the presented pipeline generates images with high fidelity but low diversity, as reflected
 266 quantitatively by comparing the perfect symmetry column with results from hot diffusion in Table 4.
 267 We attribute this to the perfect correlation between pixels of x_T sampled from the channel-wise
 268 mean Gaussian mixture model. To break the symmetry between pixels, we add a small amount of
 269 Gaussian noise (of standard deviation 0.002) to each sampled x_T . As shown in Table 4, the simple
 270 trick drastically improves the quality of generated images. We also present the qualitative results
 271 for cold diffusion using blur transformation in Figure 6, and further discuss the necessity of TACoS
 272 proposed in Algorithm 2 for generation in Appendix A.7.

Table 4: FID scores for CelebA and AFHQ datasets using hot (noise) and cold diffusion (blur transformation). Breaking the symmetry within pixels of the same channel further improves FID.

Dataset	Hot Diffusion		Cold Diffusion	
	Fixed Noise	Estimated Noise	Perfect symmetry	Broken symmetry
CelebA	59.91	23.11	97.00	49.45
AFHQ	25.62	20.59	93.05	54.68

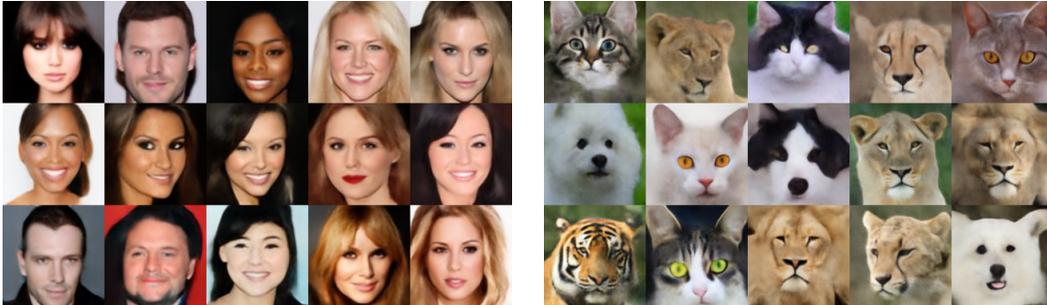


Figure 6: Examples of generated samples from 128×128 CelebA and AFHQ datasets using cold diffusion with blur transformation

273 5.3 Generation using other transformations

274 In this section, we provide a proof of concept that generation can be extended to other transformations.
 275 Specifically, we show preliminary results on inpainting, super-resolution, and *animorphosis*. Inspired
 276 by the simplicity of the degraded image distribution for the blurring routine presented in the previous
 277 section, we use degradation routines with predictable final distributions here as well.

278 To use the Gaussian mask transformation for generation, we modify the masking routine so the
 279 final degraded image is completely devoid of information. One might think a natural option is to

280 send all of the images to a completely black image x_T , but this would not allow for any diversity
 281 in generation. To get around this maximally non-injective property, we instead make the mask turn
 282 all pixels to a random, solid color. This still removes all of the information from the image, but it
 283 allows us to recover different samples from the learned distribution via Algorithm 2 by starting off
 284 with different color images. More formally, a Gaussian mask $G_t = \prod_{i=1}^t z_{\beta_i}$ is created in a similar
 285 way as discussed in the Section 4.2, but instead of multiplying it directly to the image x_0 , we create
 286 x_t as $G_t \cdot x_0 + (1 - G_t) \cdot c$, where c is an image of a randomly sampled color.

287 For super-resolution, the routine down-samples to a resolution of 2×2 , or 4 values in each channel.
 288 These degraded images can be represented as one-dimensional vectors, and their distribution is
 289 modeled using one Gaussian distribution. Using the same methods described for generation using
 290 blurring described above, we sample from this Gaussian-fitted distribution of the lower-dimensional
 291 degraded image space and pass this sampled point through the generation process trained on super-
 292 resolution data to create one output.

293 Additionally to show one can invert nearly any transformation, we include a new transformation
 294 deemed *animorphosis*, where we iteratively transform a human face from CelebA to an animal
 295 face from AFHQ. Though we chose CelebA and AFHQ for our experimentation, in principle such
 296 interpolation can be done for any two initial data distributions.

297 More formally, given an image x and a random image z sampled from the AFHQ manifold, x_t can
 298 be written as $x_t = \sqrt{\alpha_t}x + \sqrt{1 - \alpha_t}z$. Note this is essentially the same as the noising procedure,
 299 but instead of adding noise we are adding a progressively higher weighted AFHQ image. In order to
 300 sample from the learned distribution, we sample a random image of an animal and use TACoS.

301 We present results for the CelebA dataset, and hence the quantitative results in terms of FID scores
 302 for inpainting, super-resolution and *animorphosis* are 90.14, 92.91 and 48.51 respectively. We further
 303 show some qualitative samples in Figure 7, and in Figure 1.



Figure 7: Preliminary demonstration of the generative abilities of other cold diffusins on the 128×128 CelebA dataset. The top row is with *animorphosis* models, the middle row is with inpainting models, and the bottom row exhibits super-resolution models.

304 6 Conclusion

305 Existing diffusion models rely on Gaussian noise for both forward and reverse processes. In this
 306 work, we find that the random noise can be removed entirely from the diffusion model framework,
 307 and replaced with arbitrary transforms. In doing so, our generalization of diffusion models and
 308 their sampling procedures allows us to restore images afflicted by deterministic degradations such as
 309 blur, inpainting and downsampling. This framework paves the way for a more diverse landscape of
 310 diffusion models beyond the Gaussian noise paradigm. The different properties of these diffusins
 311 may prove useful for a range of applications, including image generation and beyond.

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380 **A Appendix**

381 **A.1 Deblurring**

382 For the deblurring experiments, we train the models on different datasets for 700,000 gradient steps.
383 We use the Adam Kingma and Ba [2014] optimizer with learning rate 2×10^{-5} . The training was
384 done on the batch size of 32, and we accumulate the gradients every 2 steps. Our final model is an
385 Exponential Moving Average of the trained model with decay rate 0.995 which is updated after every
386 10 gradient steps.

387 For the MNIST dataset, we blur recursively 40 times, with a discrete Gaussian kernel of size 11x11
388 and a standard deviation 7. In the case of CIFAR-10, we recursively blur with a Gaussian kernel
389 of fixed size 11x11, but at each step t , the standard deviation of the Gaussian kernel is given by
390 $0.01 * t + 0.35$. The blur routine for CelebA dataset involves blurring images with a Gaussian kernel
391 of 15x15 and the standard deviation of the Gaussian kernel grows exponentially with time t at the
392 rate of 0.01.

393 Figure 8 shows an additional nine images for each of MNIST, CIFAR-10 and CelebA. Figures 19 and
394 20 show the iterative sampling process using a deblurring model for ten example images from each
395 dataset. We further show 400 random images to demonstrate the qualitative results in the Figure 21.

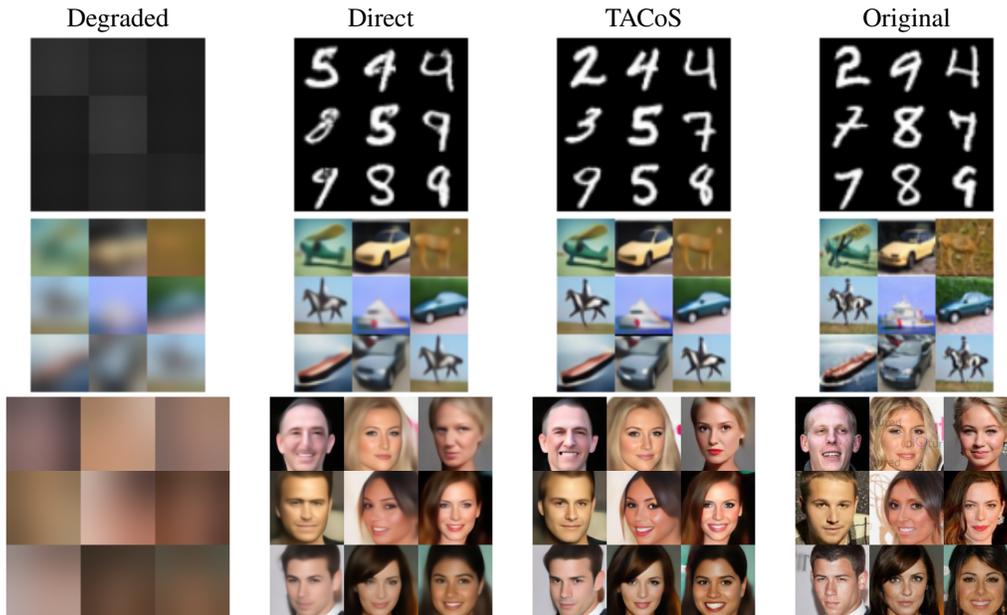


Figure 8: Additional examples from deblurring models trained on the MNIST, CIFAR-10, and CelebA datasets. **Left to right:** degraded inputs $D(x_0, T)$, direct reconstruction $R(D(x_0, T))$, sampled reconstruction with TACoS described in Algorithm 2, and original image.

396 **A.2 Inpainting**

397 For the inpainting transformation, models were trained on different datasets with 60,000 gradient
398 steps. The models were trained using Adam Kingma and Ba [2014] optimizer with learning rate
399 2×10^{-5} . We use batch size 64, and the gradients are accumulated after every 2 steps. The final model
400 is an Exponential Moving Average of the trained model with decay rate 0.995. This EMA model
401 is updated after every 10 gradient steps. For all our inpainting experiments we use a randomized
402 Gaussian mask and $T = 50$ with $\beta_1 = 1$ and $\beta_{i+1} = \beta_i + 0.1$.

403 To avoid potential leakage of information due to floating point computation of the Gaussian mask,
404 we discretize the masked image before passing it through the inpainting model. This was done by
405 rounding all pixel values to the eight most significant digits.

406 Figure 10 shows nine additional inpainting examples on each of the MNIST, CIFAR-10, and CelebA
 407 datasets. Figure 9 demonstrates an example of the iterative sampling process of an inpainting model
 408 for one image in each dataset.

409 A.3 Super-Resolution

410 We train the super-resolution model per Section 3.1 for 700,000 iterations. We use the Adam Kingma
 411 and Ba [2014] optimizer with learning rate 2×10^{-5} . The batch size is 32, and we accumulate the
 412 gradients every 2 steps. Our final model is an Exponential Moving Average of the trained model with
 413 decay rate 0.995. We update the EMA model every 10 gradient steps.

414 The number of time-steps depends on the size of the input image and the final image. For MNIST and
 415 for CIFAR10, the number of time steps is 3, as it takes three steps of halving the resolution to reduce
 416 the initial image down to 4×4 . For CelebA, the number of time steps is 6 to reduce the initial image
 417 down to 2×2 . For CIFAR10, we apply random crop and random horizontal flip for regularization.

418 Figure 12 shows an additional nine super-resolution examples on each of the MNIST, CIFAR-10, and
 419 CelebA datasets. Figure 11 shows one example of the progressive increase in resolution achieved
 420 with the sampling process using a super-resolution model for each dataset.

421 A.4 Colorization

422 Here we provide results for the additional task of colorization. Starting with the original RGB-
 423 image x_0 , we realize colorization by iteratively desaturating for T steps until the final image x_T
 424 is a fully gray-scale image. We use a series of three-channel 1×1 convolution filters $\mathbf{z}(\alpha) =$
 425 $\{z^1(\alpha), z^2(\alpha), z^3(\alpha)\}$ with the form

$$z^1(\alpha) = \alpha \begin{pmatrix} \frac{1}{3} & \frac{1}{3} & \frac{1}{3} \end{pmatrix} + (1 - \alpha) \begin{pmatrix} 1 & 0 & 0 \end{pmatrix}$$

$$z^2(\alpha) = \alpha \begin{pmatrix} \frac{1}{3} & \frac{1}{3} & \frac{1}{3} \end{pmatrix} + (1 - \alpha) \begin{pmatrix} 0 & 1 & 0 \end{pmatrix}$$

$$z^3(\alpha) = \alpha \begin{pmatrix} \frac{1}{3} & \frac{1}{3} & \frac{1}{3} \end{pmatrix} + (1 - \alpha) \begin{pmatrix} 0 & 0 & 1 \end{pmatrix}$$

428 and obtain $D(x, t) = \mathbf{z}(\alpha_t) * x$ via a schedule defined as $\alpha_1, \dots, \alpha_t$ for each respective step. Notice
 429 that a gray image is obtained when $x_T = \mathbf{z}(1) * x_0$.

430 We can tune the ratio α_t to control the amount of information removed in each step. For our
 431 experiment, we schedule the ratio such that for every t we have

$$x_t = \mathbf{z}(\alpha_t) * \dots * \mathbf{z}(\alpha_1) * x_0 = \mathbf{z}\left(\frac{t}{T}\right) * x_0.$$

432 This schedule ensures that color information lost between steps is smaller in earlier stage of the
 433 diffusion and becomes larger as t increases.

434 We train the models on different datasets for 700,000 gradient steps. We use Adam Kingma and Ba
 435 [2014] optimizer with learning rate 2×10^{-5} . We use batch size 32, and we accumulate the gradients
 436 every 2 steps. Our final model is an exponential moving average of the trained model with decay rate
 437 0.995. We update the EMA model every 10 gradient steps. For CIFAR-10 we use $T = 50$ and for
 438 CelebA we use $T = 20$.

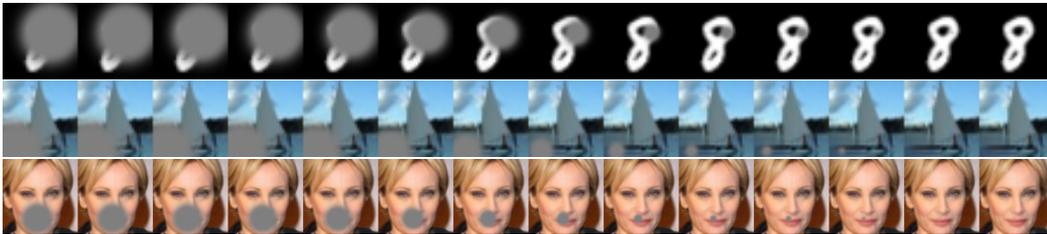


Figure 9: Progressive inpainting of selected masked MNIST, CIFAR-10, and CelebA images.

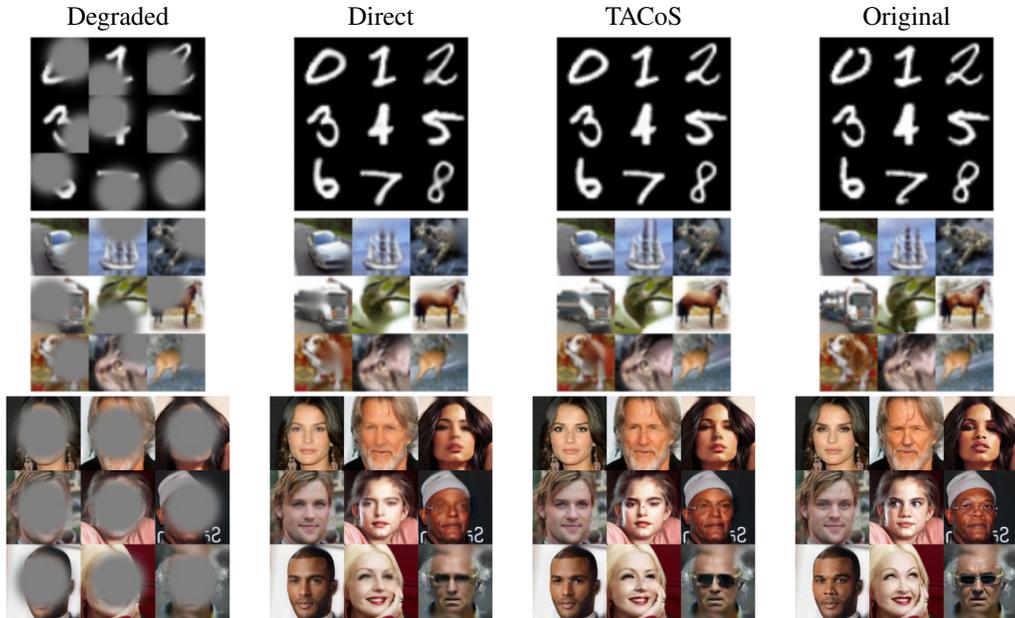


Figure 10: Additional examples from inpainting models trained on the MNIST, CIFAR-10, and CelebA datasets. **Left to right:** degraded inputs $D(x_0, T)$, direct reconstruction $R(D(x_0, T))$, sampled reconstruction with TACoS described in Algorithm 2, and original image.



Figure 11: Progressive upsampling of selected downsampled MNIST, CIFAR-10, and CelebA images. The original image is at the left for each of these progressive upsamplings.

439 We illustrate our recolorization results in Figure 13. We present testing examples, as well as their
 440 grey scale images, from all the datasets, and compare the recolorization results with the original
 441 images. The recolored images feature correct color separation between different regions, and feature
 442 various and yet semantically correct colorization of objects. Our sampling technique still yields
 443 minor differences in comparison to the direct reconstruction, although the change is not visually
 444 apparent. We attribute this to the shape restriction of colorization task, as human perception is rather
 445 insensitive to minor color change. We also provide quantitative measurement for the effectiveness
 446 of our recolorization results in terms of different similarity metrics, and summarize the results in
 447 Table 5.

Table 5: Quantitative metrics for quality of image reconstruction using recolorization models for all three channel datasets.

Dataset	Degraded Image			Reconstruction		
	FID	SSIM	RMSE	FID	SSIM	RMSE
CIFAR-10	97.39	0.937	0.078	45.74	0.942	0.069
CelebA	41.20	0.942	0.089	17.50	0.973	0.042

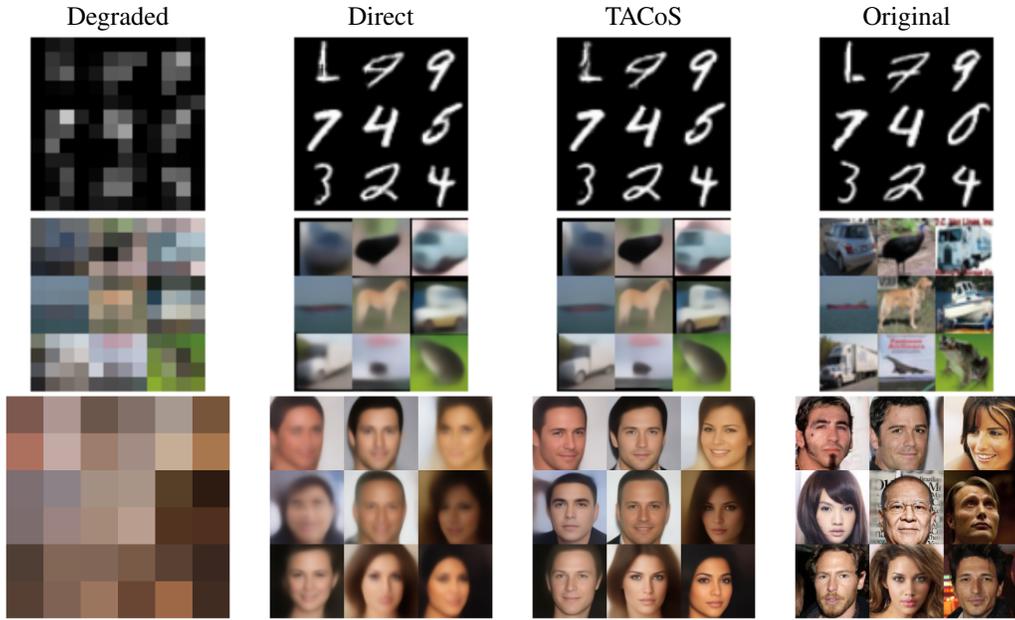


Figure 12: Additional examples from super-resolution models trained on the MNIST, CIFAR-10, and CelebA datasets. **Left to right:** degraded inputs $D(x_0, T)$, direct reconstruction $R(D(x_0, T))$, sampled reconstruction with TACoS described in Algorithm 2, and original image.

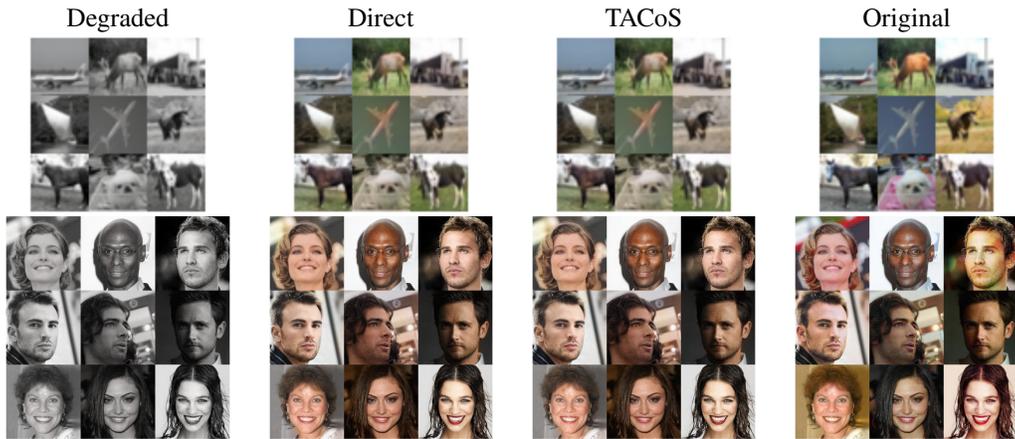


Figure 13: Recolorization models trained on the CIFAR-10 and CelebA datasets. **Left to right:** degraded inputs $D(x_0, T)$, direct reconstruction $R(D(x_0, T))$, sampled reconstruction with TACoS described in Algorithm 2, and original image.

448 A.5 Image Snow

449 Here we provide results for the additional task of snowification, which is a direct adaptation of
 450 the official implementation of ImageNet-C snowification process Hendrycks and Dietterich [2019].
 451 To determine the snow pattern of a given image $x_0 \in \mathbb{R}^{C \times H \times W}$, we first construct a seed matrix
 452 $S_A \in \mathbb{R}^{H \times W}$ where each entry is sampled from a Gaussian distribution $N(\mu, \sigma)$. The upper-left
 453 corner of S_A is then zoomed into another matrix $S_B \in \mathbb{R}^{H \times W}$ with spline interpolation. Next, we
 454 create a new matrix S_C by filtering each value of S_B with a given threshold c_1 as

$$S_C[i][j] = \begin{cases} 0, & S_B[i][j] \leq c_1 \\ S_B[i][j], & S_B[i][j] > c_1 \end{cases}$$

455 and clip each entry of S_C into the range $[0, 1]$. We then convolve S_C using a motion blur kernel with
 456 standard deviation c_2 to create the snow pattern S and its up-side-down rotation S' . The direction of
 457 the motional blur kernel is randomly chosen as either vertical or horizontal. The final snow image is
 458 created by again clipping each value of $x_0 + S + S'$ into the range $[0, 1]$. For simplicity, we abstract
 459 the process as a function $h(x_0, S_A, c_0, c_1)$.

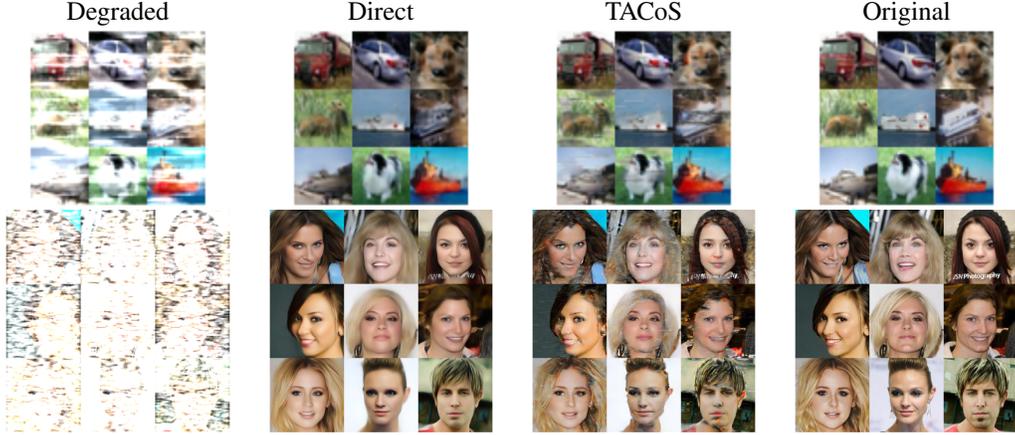


Figure 14: Additional examples from *Desnowification* models trained on the CIFAR-10 and CelebA datasets. **Left to right:** degraded inputs $D(x_0, T)$, direct reconstruction $R(D(x_0, T))$, sampled reconstruction with TACoS described in Algorithm 2, and original image.

460 To create a series of T images with increasing snowification, we linearly interpolate c_0 and c_1 between
 461 $[c_0^{\text{start}}, c_0^{\text{end}}]$ and $[c_1^{\text{start}}, c_1^{\text{end}}]$ respectively, to create $c_0(t)$ and $c_1(t)$, $t = 1, \dots, T$. Then for each x_0 , a
 462 seed matrix S_x is sampled, the motion blur direction is randomized, and we construct each related
 463 x_t by $x_t = h(x_0, S_x, c_0(t), c_1(t))$. Visually, $c_0(t)$ dictates the severity of the snow, while $c_1(t)$
 464 determines how “windy” the snowified image seems.

465 For both CIFAR-10 and Celeb-A, we use the same Gaussian distribution with parameters $\mu = 0.55$
 466 and $\sigma = 0.3$ to generate the seed matrix. For CIFAR-10, we choose $c_0^{\text{start}} = 1.15$, $c_0^{\text{end}} = 0.7$,
 467 $c_1^{\text{start}} = 0.05$ and $c_1^{\text{end}} = 16$, which generates a visually lighter snow. For Celeb-A, we choose
 468 $c_0^{\text{start}} = 1.15$, $c_0^{\text{end}} = 0.55$, $c_1^{\text{start}} = 0.05$ and $c_1^{\text{end}} = 20$, which generates a visually heavier snow.

469 We train the models on different datasets for 700,000 gradient steps. We use Adam Kingma and Ba
 470 [2014] optimizer with learning rate 2×10^{-5} . We use batch size 32, and we accumulate the gradients
 471 every 2 steps. Our final model is an exponential moving average of the trained model with decay
 472 rate 0.995. We update the EMA model every 10 gradient steps. For CIFAR-10 we use $T = 200$ and
 473 for CelebA we use $T = 200$. We note that the seed matrix is resampled for each individual training
 474 batch, and hence the snow pattern varies across the training stage.

475 A.6 Generation using noise : Further Details

476 Here we show the equivalence between the sampling method proposed in Algorithm 2 and the
 477 deterministic sampling in DDIM Song et al. [2021a]. Given the image x_t at step t , we have the
 478 restored clean image \hat{x}_0 from the diffusion model. Hence given the estimated \hat{x}_0 and x_t , we can
 479 estimate the noise $z(x_t, t)$ (or \hat{z}) as

$$z(x_t, t) = \frac{x_t - \sqrt{\alpha_t} \hat{x}_0}{\sqrt{1 - \alpha_t}},$$

480 Thus, the $D(\hat{x}_0, t)$ and $D(\hat{x}_0, t - 1)$ can be written as

$$D(\hat{x}_0, t) = \sqrt{\alpha_t} \hat{x}_0 + \sqrt{1 - \alpha_t} \hat{z},$$

481

$$D(\hat{x}_0, t - 1) = \sqrt{\alpha_{t-1}} \hat{x}_0 + \sqrt{1 - \alpha_{t-1}} \hat{z},$$

482 using which the sampling process in Algorithm 2 to estimate x_{t-1} can be written as,

$$\begin{aligned}
x_{t-1} &= x_t - D(\hat{x}_0, t) + D(\hat{x}_0, t-1) \\
&= x_t - (\sqrt{\alpha_t}\hat{x}_0 + \sqrt{1-\alpha_t}\hat{z}) + (\sqrt{\alpha_{t-1}}\hat{x}_0 + \sqrt{1-\alpha_{t-1}}\hat{z}) \\
&= \sqrt{\alpha_{t-1}}\hat{x}_0 + \sqrt{1-\alpha_{t-1}}\hat{z}
\end{aligned}
\tag{1}$$

483 which is same as the sampling method as described in Song et al. [2021a]. The only difference from
484 the original Song et al. [2021a] is the order for estimating \hat{x}_0 and \hat{z} . The original Song et al. [2021a]
485 paper estimated \hat{z} first and then used this to predict clean image \hat{x}_0 , while we first predict the clean
486 image \hat{x}_0 and then estimate the noise \hat{z} .

487 A.7 Generation using blur transformation: Further Details

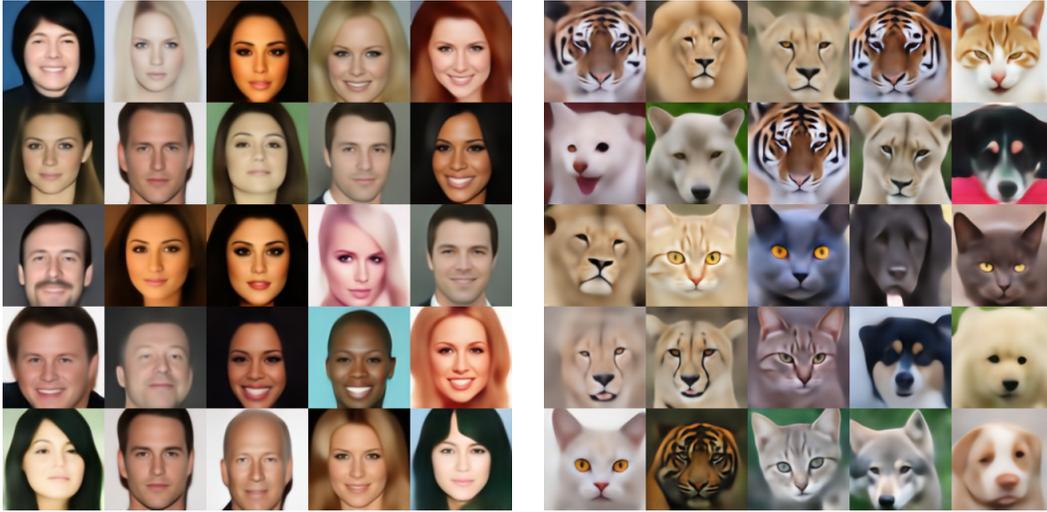


Figure 15: Examples of generated samples from 128×128 CelebA and AFHQ datasets using Method 2 with perfect symmetry.

488 Figure 15 shows that generation without breaking any symmetry within each channel is quite
489 promising as well.

490 **Necessity of Algorithm 2:** In the case of unconditional generation, we observe a marked superiority
491 in quality of the sampled reconstruction using Algorithm 2 over any other method considered. For
492 example, in the broken symmetry case, the FID of the directly reconstructed images is 257.69 for
493 CelebA and 214.24 for AFHQ, which are far worse than the scores of 49.45 and 54.68 from Table 4.
494 In Figure 17, we also give a qualitative comparison of this difference. We can also clearly see from
495 Figure 18 that Algorithm 1, the method used in Song et al. [2021b] and Ho et al. [2020], completely
496 fails to produce an image close to the target data distribution.

497 A.8 Algorithm 1 is same as DDIM/DDPM sampling

498 The sampling method proposed in Song et al. [2021a] in it’s equation 12 is given as

$$x_{t-1} = \sqrt{\alpha_{t-1}} \cdot \text{“predicted } x_0 \text{”} + \sqrt{1 - \alpha_{t-1} - \sigma_t^2} \epsilon_\theta(x_t) + \sigma_t \epsilon_t$$

499 where $\epsilon_\theta(x_t)$ is the noise predicted by the diffusion model given x_t and t . The term “predicted x_0 ” or
500 \hat{x}_0 can be computed directly given x_t and $\epsilon_\theta(x_t)$ as

$$\hat{x}_0 = \frac{x_t - \sqrt{1 - \alpha_t} \epsilon_\theta(x_t)}{\sqrt{\alpha_t}},$$

501 Hence using \hat{z} instead of $\epsilon_\theta(x_t)$ and \hat{x}_0 to indicate predicted clean image, we have

$$x_{t-1} = \sqrt{\alpha_{t-1}} \cdot \hat{x}_0 + \sqrt{1 - \alpha_{t-1} - \sigma_t^2} \hat{z} + \sigma_t \epsilon_t$$

502 Thus, the sampling step can interpreted as follows: At each step t , we start with a noisy image x_t and
 503 use the diffusion model to estimate the clean image \hat{x}_0 and the noise \hat{z} that was added to this clean
 504 image \hat{x}_0 to get the noisy image x_t . In order to move to lesser noisy image x_{t-1} , one ‘‘adds back’’
 505 lesser noise to the the ‘‘predicted clean image’’ \hat{x}_0 . Now one can add back noise in 2 ways, either the
 506 noise which was added to the clean image \hat{x}_0 which is \hat{z} or sample a new uncorrelated noise ϵ_t . Infact
 507 both of these noise can be added using σ_t as the hyperparameter that weighs the amount of each noise
 508 added. This σ_t is placed in the equation such that for any choice of σ_t , the standard deviation of noise
 509 added back is $\sqrt{1 - \alpha_{t-1}}$. For $\sigma_t = 0$, we only add back the estimated noise \hat{z} and no uncorrelated
 510 noise ϵ_t which is infact the DDIM sampling. While for $\sigma_t = \sqrt{(1 - \alpha_{t-1})/(1 - \alpha_t)} \sqrt{1 - \alpha_t/\alpha_{t-1}}$
 511 we get the sampling method described in DDPM.

512 Nevertheless, for any choice of σ_t , the sampling method involves a denoising operation which is
 513 shown as $R(x_s, s)$ in Algorithm 1 and adding back noise shown as $x_{s-1} = D(\hat{x}_0, s-1)$ in Algorithm
 514 1. The only difference between different sampling methods explained in DDPM or DDIM is how one
 515 degrades the image back.

516 A.9 Discussion on Algorithm 2 produces lesser error than Algorithm 1

517 Consider the case where the degradation is a simple blur operator that removes one frequency from
 518 an image every time t increases by one. We can write $X = x_0 + x_1 \dots x_{T-1} + x_T$, where each
 519 x_i is the Fourier mode of X representing frequency $T - i$. The degradation operator for this blur
 520 is $D(X, t) = x_t + \dots x_{T-1} + x_T = \sum_{i=t}^T x_i$, and when $t = T$ the signal has been blurred into a
 521 constant vector.

522 Suppose we are performing the reverse process, and we begin at step t with $X_t = x_t + \dots x_T$. Now
 523 in order to go from X_t to X_{t-1} , we first use the diffusion model to predict $\hat{X} = R(X_t, t)$. This \hat{X}
 524 has an analogous Fourier expansion $\hat{X} = \sum_{n=0}^T \hat{x}_n$, where \hat{x}_0 is the highest frequency mode and \hat{x}_T
 525 is the DC mode. At step t , the error E_t can be defined as $\|X_{t-1} - \hat{X}_{t-1}\|$.

526 For the Algorithm 1, in which X_{t-1} is given by $\hat{X}_{t-1} = D(\hat{X}, t-1)$, we can expand the error as
 527 follows

$$\begin{aligned} E_t^2 &= \|X_{t-1} - \hat{X}_{t-1}\|^2 \\ &= \left\| \sum_{i=t-1}^T x_i - \sum_{i=t-1}^T \hat{x}_i \right\|^2 \\ &= \left\| \sum_{i=t}^T (x_i - \hat{x}_i) \right\|^2 \\ &= \sum_{i=t-1}^T \|x_i - \hat{x}_i\|^2 \end{aligned}$$

528 For the Algorithm 2, in which X_{t-1} is given by $\hat{X}_{t-1} = X_t - D(\hat{X}, t) + D(\hat{X}, t-1)$, we can
 529 expand the error as follows

$$\begin{aligned} E_t^2 &= \|X_{t-1} - (X_t - D(\hat{X}, t) + D(\hat{X}, t-1))\|^2 \\ &= \|(D(\hat{X}, t) - D(\hat{X}, t-1)) - (X_t - X_{t-1})\|^2 \\ &= \left\| \left(\sum_{i=t}^T \hat{x}_i - \sum_{i=t-1}^T \hat{x}_i \right) - \left(\sum_{i=t}^T x_i - \sum_{i=t-1}^T x_i \right) \right\|^2 \\ &= \|\hat{x}_{t-1} - x_{t-1}\|^2 \end{aligned}$$

530 Hence, we can see that the error incurred at step t using algorithm 1 is $\sum_{i=t-1}^T \|x_i - \hat{x}_i\|^2$ which is
 531 clearly greater than or equal to error incurred by using the algorithm 2 which is $\|\hat{x}_{t-1} - x_{t-1}\|^2$. We

532 now further break down the analysis based on different scenarios depending on the quality of the
533 reconstruction operator R :

534

535 **1. R is perfect.**

536 In this scenario, both Algorithm 1 and Algorithm 2 are indistinguishable, and the error is 0
537 for both.

538 **2. R is imperfect only for x_{t-1} and reconstructs other signals perfectly.**

539 In this scenario, both Algorithm 1 and Algorithm 2 incur the same error, which is $(x_{t-1} -$
540 $\hat{x}_{t-1})^2$.

541 **3. R is imperfect for more than one frequency**

542 This is the realistic scenario. In this case, the error incurred by Algorithm 1 is strictly greater
543 than Algorithm 2.

544 Thus we prove that for the realistic scenario i.e. when R is not a perfect reconstruction operator, the
545 error incurred using Algorithm 1 is always greater than the Algorithm 2.

546 **A.10 Empirical comparisons between Algorithm 2 and Algorithm 1**

547 For all the degradations in Section 4, we compare to "Direct Sampling" which is in fact the "one step"
548 reconstruction of the input to Algorithm 2. One of the main contributions of our work is Algorithm
549 2, which outperforms Algorithm 1 across all degradations. Infact the Algorithm 1 is worse than the
550 one-step generation as well. To demonstrate this concretely we present the FID results in 6, where
551 we can clearly see that FID scores for Algorithm 1 are worse than both the one-step and Algorithm 2.
552 We also show qualitatively how bad the Algorithm 1 is in Figures 16 and 18. In fact, for the case
553 of the celebA dataset, Algorithm 1 fails drastically, while for CIFAR-10 we can see high-frequency
554 signals present in the image generation. Hence, we chose our baseline to be direct generation instead
555 of Algorithm 1.

Table 6: FIDs for blur degradation for Algorithm 1, Algorithm 2 and Direct Reconstruction. This table demonstrates that Algorithm 1 is worse than both the one-step generation and Algorithm 2

Dataset	Direct Generation	Algorithm 1	Algorithm 2
MNIST	5.10	8.24	4.69
CIFAR-10	83.69	97.89	80.08
CelebA	36.37	299.61	26.14

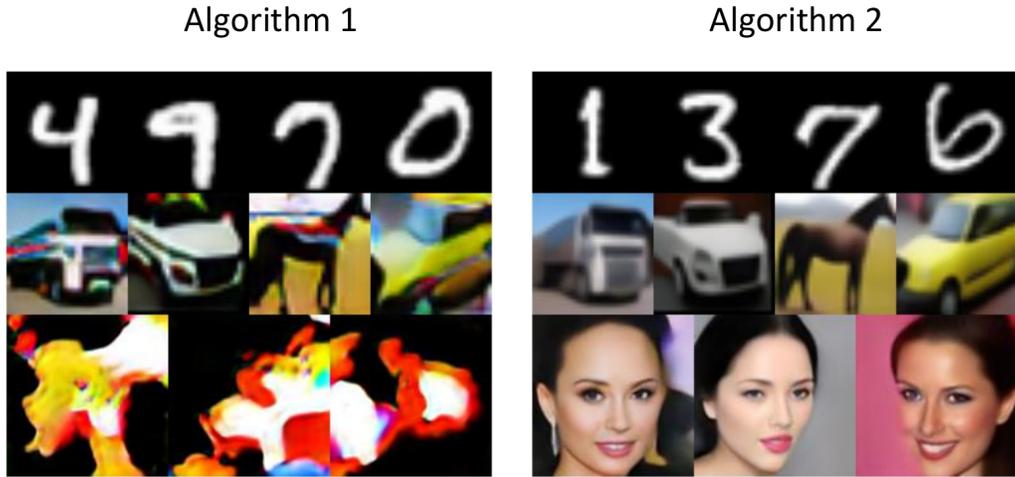


Figure 16: Comparison of Algorithm 1 and Algorithm 2. We demonstrate that Algorithm 1 performs much worse than Algorithm 1 and fails completely for CelebA dataset.

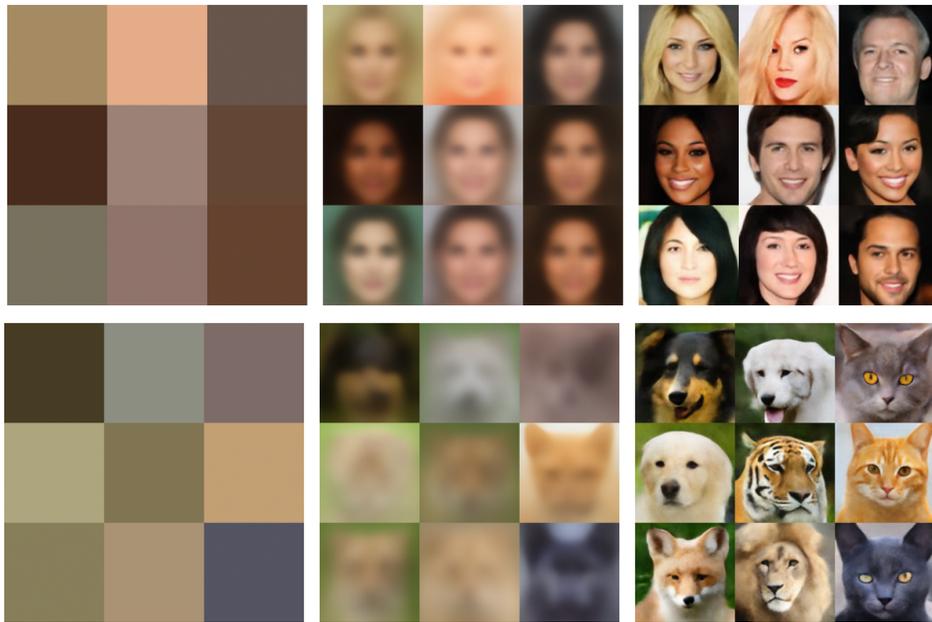


Figure 17: Comparison of direct reconstruction with sampling using TACoS described in Algorithm 2 for generation with blur transformation and broken symmetry. Left-hand column is the initial cold images generated using the simple Gaussian model. Middle column has images generated in one step (*i.e.* direct reconstruction). Right-hand column are the images sampled with TACoS described in Algorithm 2. We present results for both CelebA (top) and AFHQ (bottom) with resolution 128×128 .

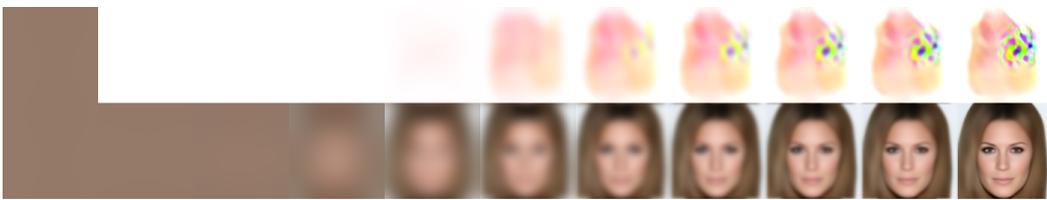


Figure 18: Comparison of Algorithm 1 (top row) and Algorithm 2 (bottom row) for generation with Method 2 and broken symmetry on 128×128 CelebA dataset. We demonstrate that Algorithm 1 fails completely to generate a new image.

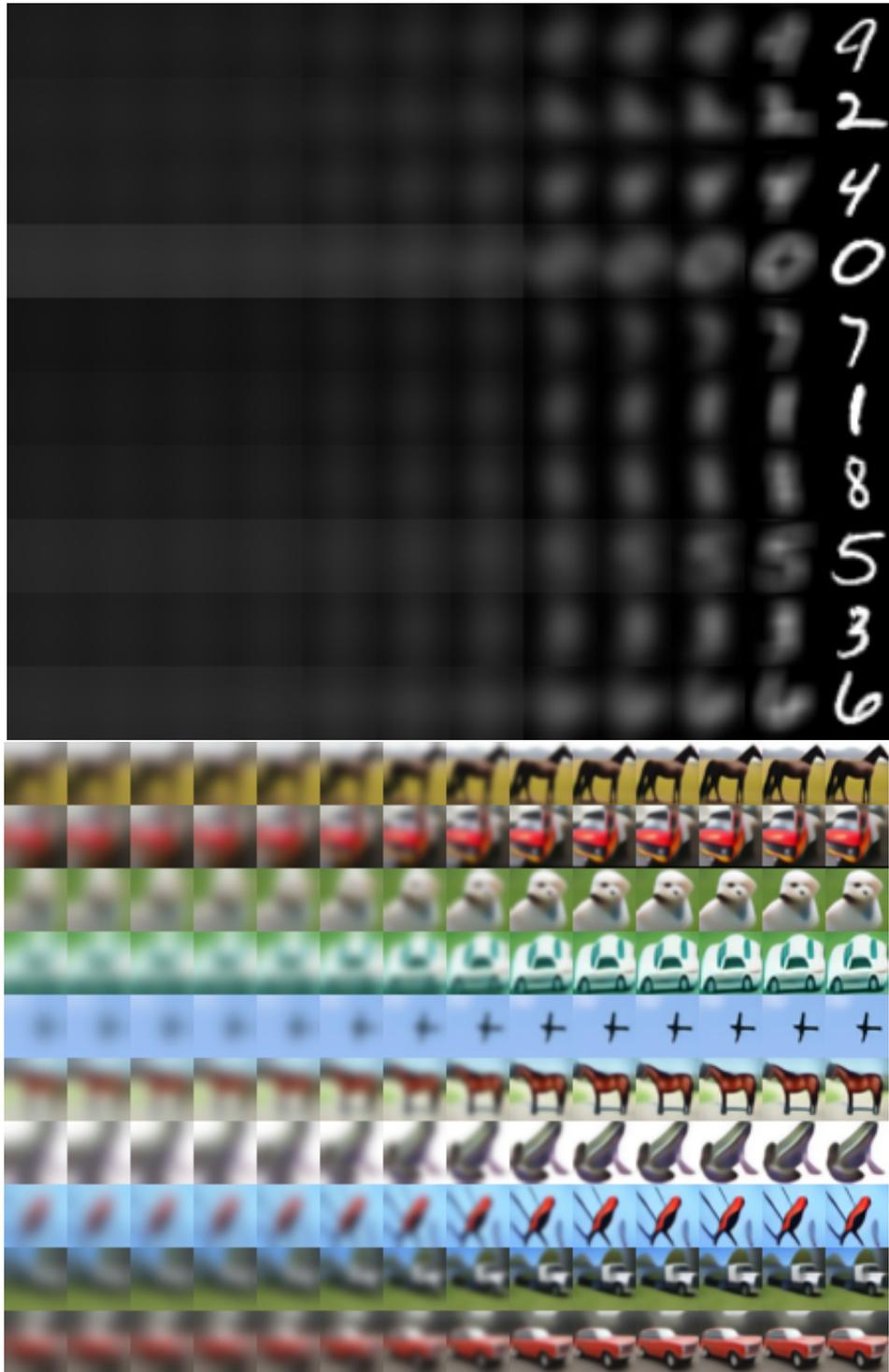


Figure 19: Progressive deblurring of selected blurred MNIST and CIFAR-10 images.

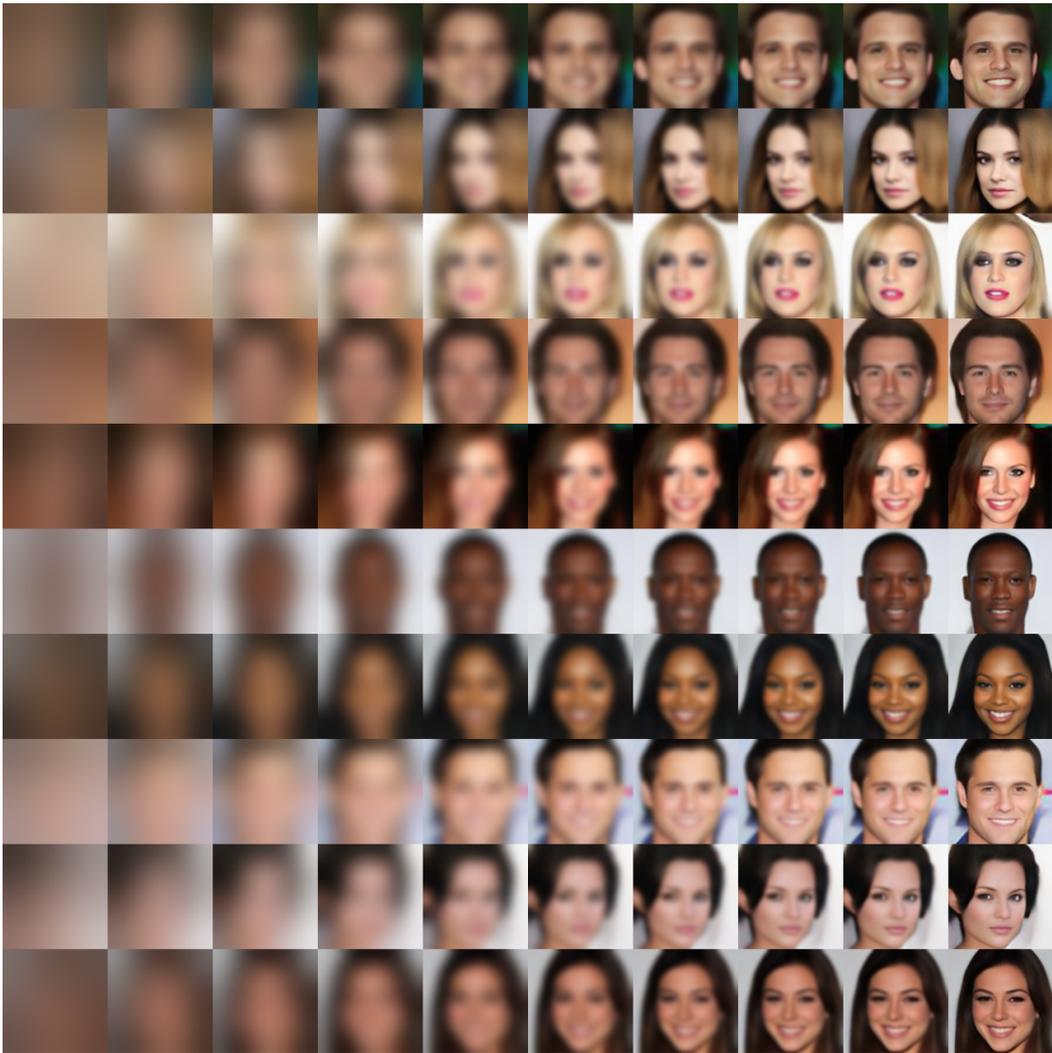


Figure 20: Progressive deblurring of selected blurred CelebA images.



Figure 21: Deblurred Cifar10 images